

安定なヒッグスボソンの コライダー実験における シグナル

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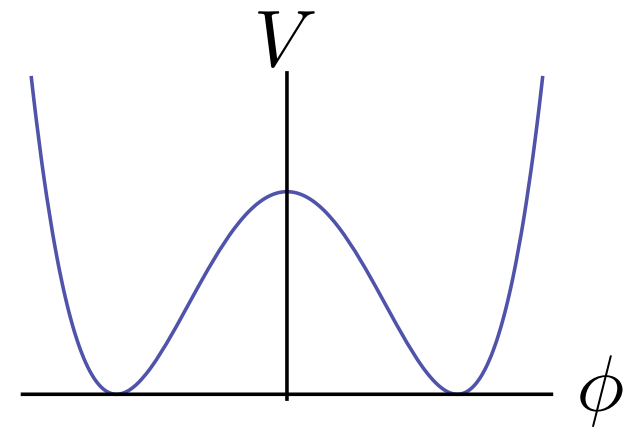
岡山大学津島キャンパス, 2010/03/22

Introduction

Electro-Weak Symmetry Breaking

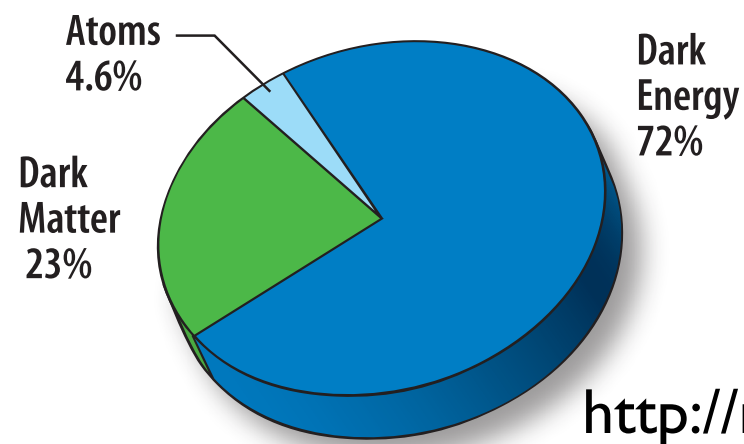
Higgs mechanism: Not seen yet.

Naturalness and hierarchy problem: New physics?



Dark Matter

WIMP?



<http://map.gsfc.nasa.gov/>

A possible solution: **Dark Higgs scenario**

Stable Higgs in gauge-Higgs unification

Model

Hosotani, Oda, Ohnuma, Sakamura, PRD78,096002(2008).

$SO(5) \times U(1)$ in 5D warped space-time.

EWSB by Hosotani mechanism.

4D Higgs field: Wilson line phase,

$$\hat{\theta}_H(x) = \theta_H + \frac{H(x)}{f_H} . \quad f_H \simeq 246 \text{ GeV}$$

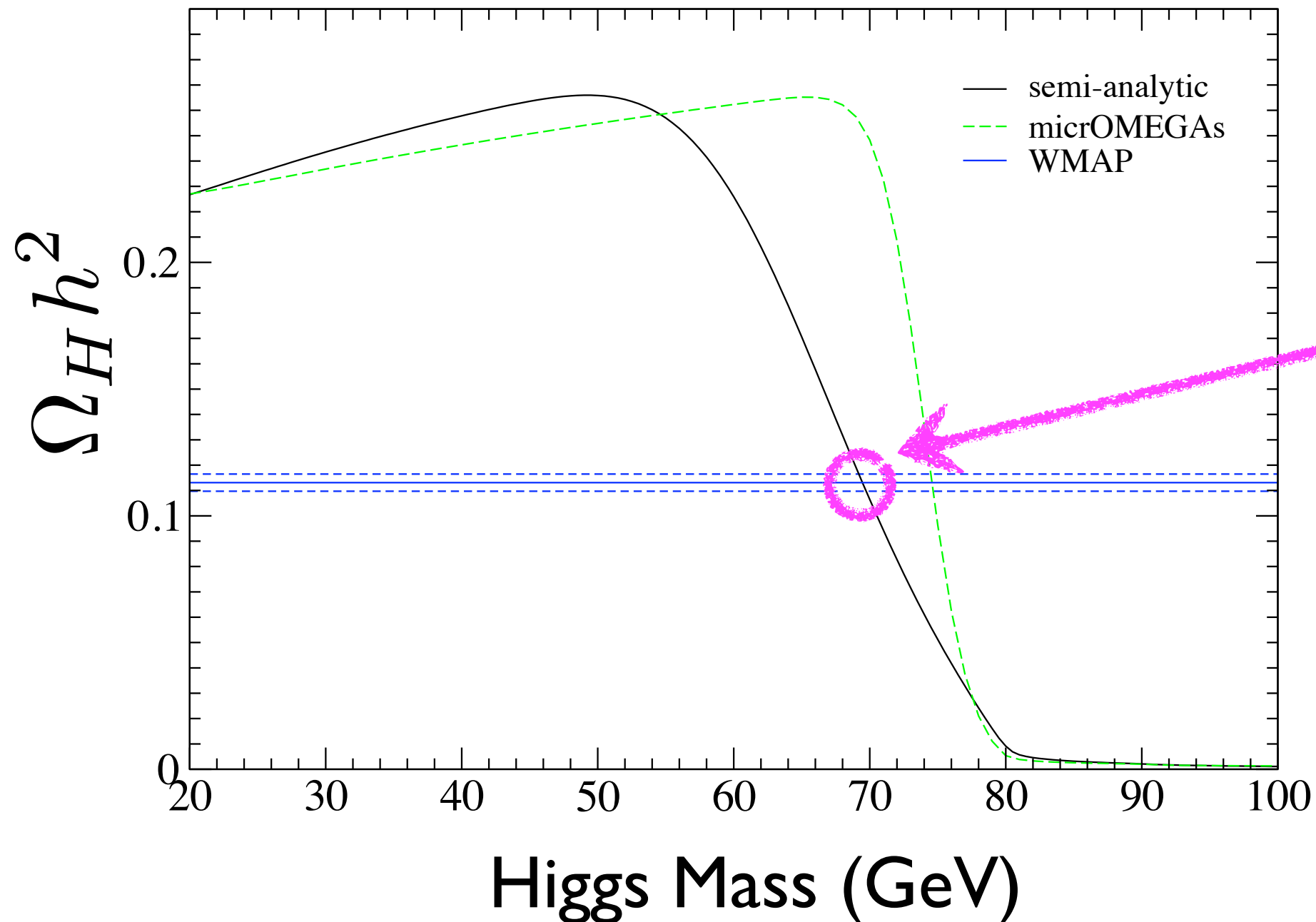


A new dynamical parity, **H-parity**,

$$H(x) \rightarrow -H(x) .$$

Higgs is STABLE!

Relic Abundance



$m_H \sim 70 \text{ GeV}$
favored.

Collider signals

Effective interactions

Integrating out KK modes,

$$\begin{aligned}\mathcal{L}_{\text{int}} = & -\frac{m_W^2}{f_H^2} H^2 W^{+\mu} W_{\mu}^{-} - \frac{m_Z^2}{2f_H^2} H^2 Z^{\mu} Z_{\mu} \\ & + \sum_f \frac{m_f}{2f_H^2} H^2 \bar{f} f + \cdots .\end{aligned}$$

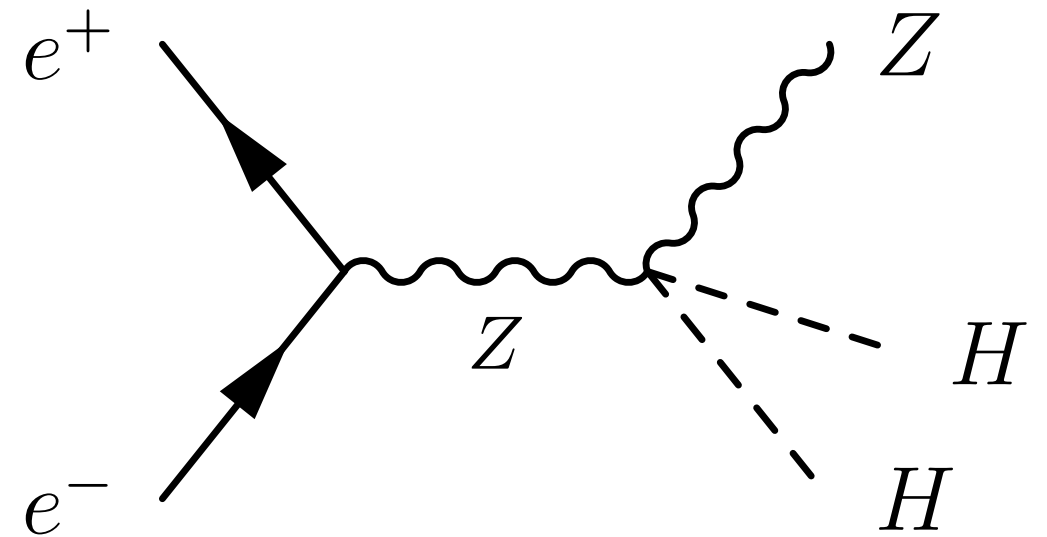
No odd powers of H .

Two missing Higgs bosons

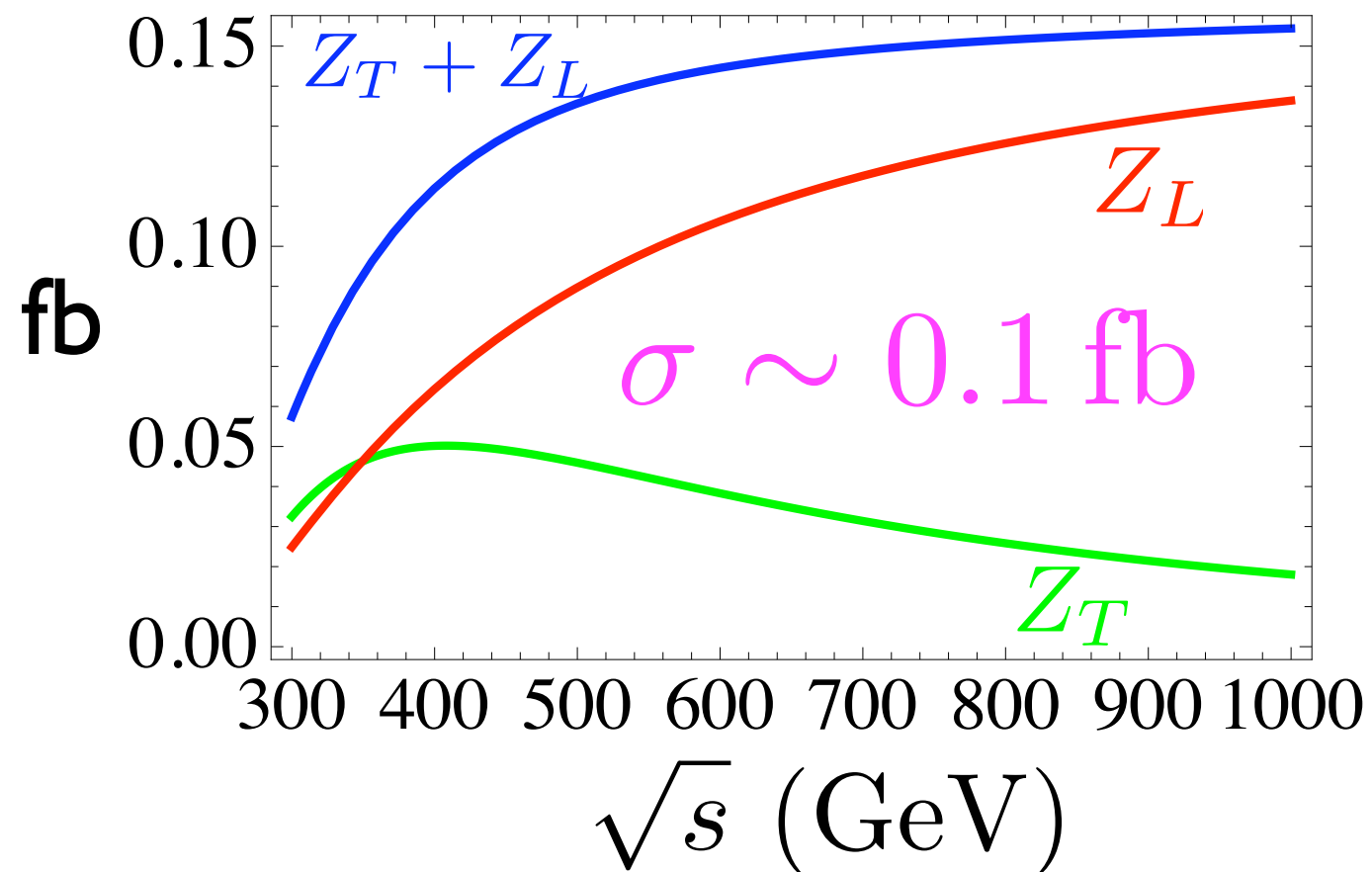
Linear Collider

Signal: $e^+e^- \rightarrow ZHH$

H's are missing.



total cross section for $m_H = 70 \text{ GeV}$



Z_L violates the unitarity
unless $s/m_{KK}^2 \ll 1$.

$m_{KK} \sim 1 \text{ TeV}$

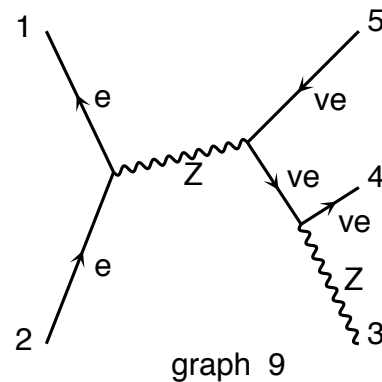
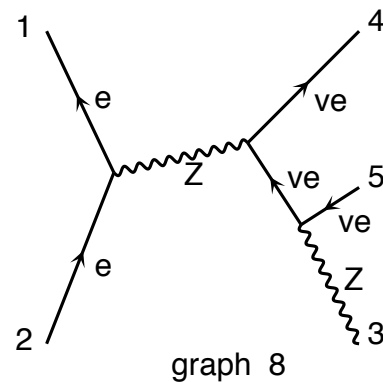
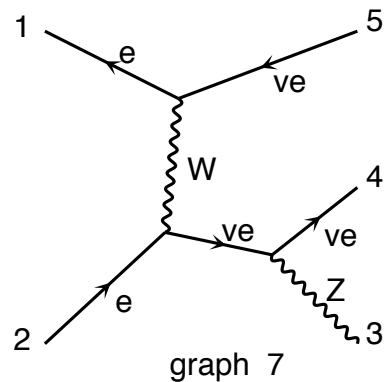
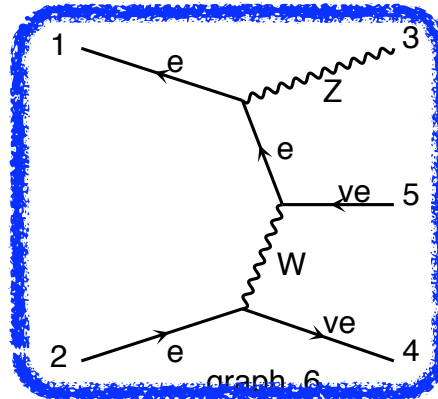
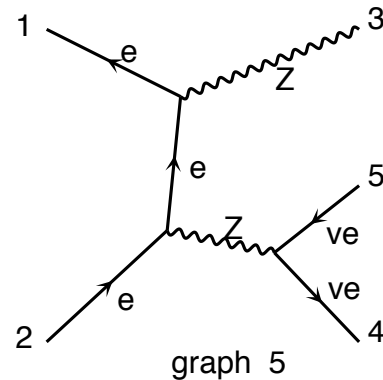
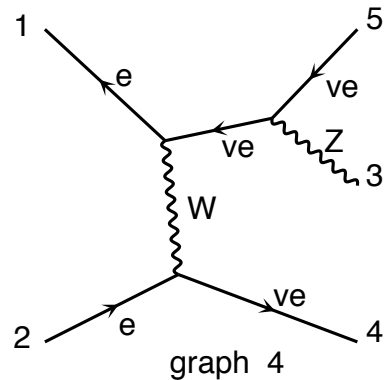
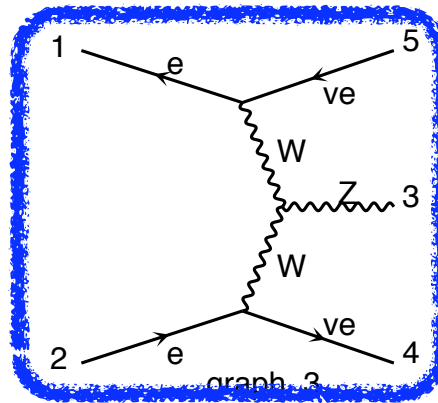
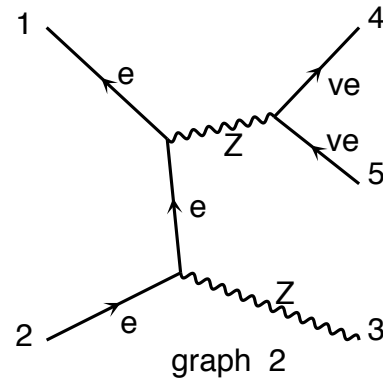
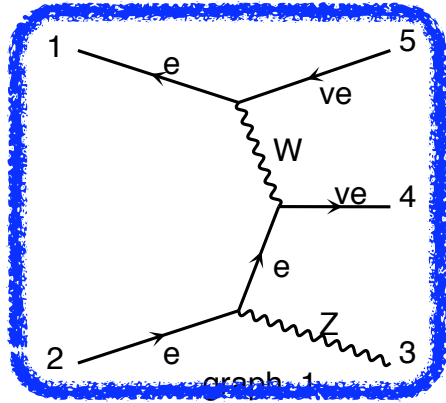
$\sqrt{s} = 500 \text{ GeV}$

in the following.

LC background

$$e^+e^- \rightarrow Z\nu\bar{\nu}$$

Diagrams by MadGraph



BG cross section with

$$M_{\text{miss}} \geq 120 \text{ GeV}$$

$$\sigma_{\text{BG}} \simeq 311 \text{ fb}$$

Need polarizations!

beams and Z

LC with polarizations

Ideal case: $e_L^+ e_R^- \rightarrow Z_L H H, Z_L \nu \bar{\nu}$

$$\sigma_{\text{signal}} \simeq 0.12 \text{ fb} \quad \textbf{vs} \quad \sigma_{\text{BG}} \simeq 0.42 \text{ fb}$$

$|\cos \theta| < 0.6$ is applied.

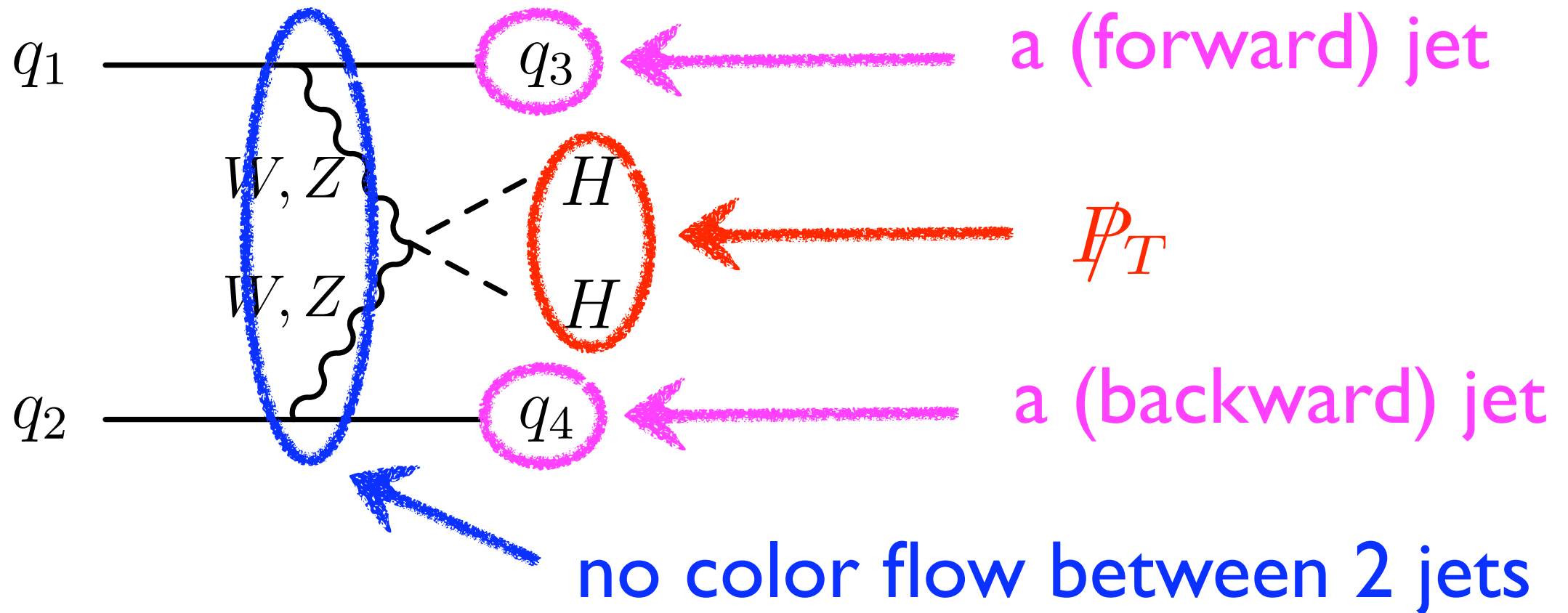
Significance: $\mathcal{S} \equiv \frac{N_{\text{signal}}}{\sqrt{N_{\text{signal}} + N_{\text{BG}}}}$

$$\mathcal{S} = 1.4 \sqrt{L / 100 \text{ fb}^{-1}}$$

A few (or more) ab^{-1} is required!

LHC

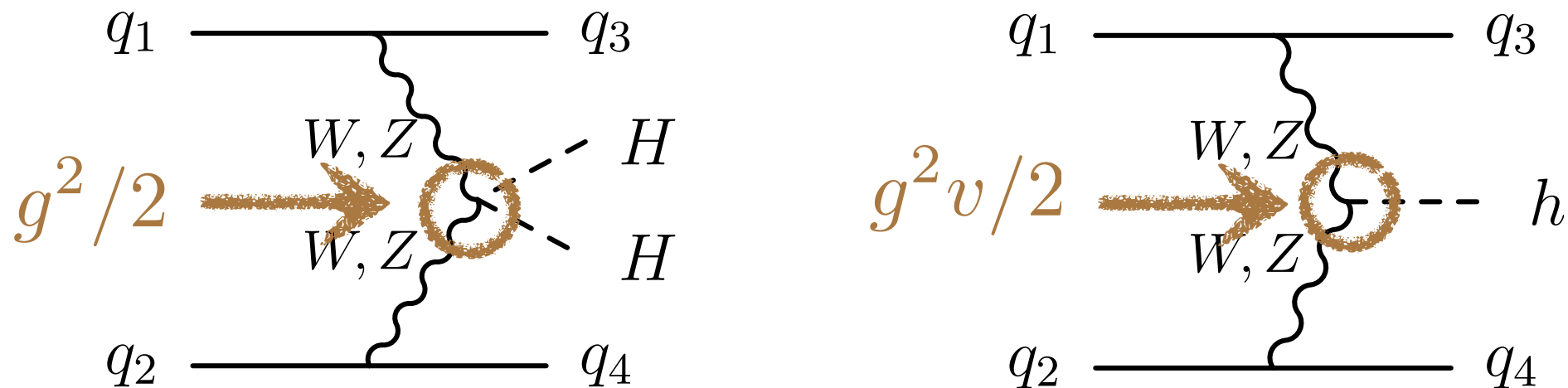
Signal: Weak boson fusion



Background: Wjj , Zjj , jjj

→ Similar as invisible Higgs search

Signal cross section at LHC



$$\frac{d\sigma_{HH}}{dm_{HH}^2} = \frac{\bar{\beta}_f}{32\pi^2 v^2} \sigma_h \big|_{m_h^2 = m_{HH}^2}$$



$$\sigma_{HH} \sim 1.5 \text{ fb}$$

$$\sigma_{BG} \simeq 167 \text{ fb}$$

$$\mathcal{S} \sim 1.2 \sqrt{L/100 \text{ fb}^{-1}}$$

in the SM

Éboli, Zeppenfeld

$$p_T^j > 40 \text{ GeV}, \quad |\eta_j| < 5.0,$$

$$|\eta_{j1} - \eta_{j2}| > 4.4, \quad \eta_{j1} \cdot \eta_{j2} < 0,$$

$$\not{p}_T > 100 \text{ GeV}.$$

$$M_{jj} > 1200 \text{ GeV}, \quad \phi_{jj} < 1.$$

Summary

- ★ Stable Higgs in gauge-Higgs unification is a viable candidate of dark matter.

Dark Higgs scenario

- ★ $m_H \sim 70 \text{ GeV}$ is predicted.
- ★ We need **a few ab^{-1}** or more.
both for LHC and LC.
- ★ Signals in KK mode production should be studied.
 $m_{\text{KK}} \lesssim 1 \text{ TeV}$

Backup Slides

Spin-Independent Cross Section

CDMS II

arXiv:0912.3592

Local DM density

$$\rho_0 = 0.3 \text{ GeV}/\text{cm}^3$$

assumed in exps.

For $m_H = 70 \text{ GeV}$

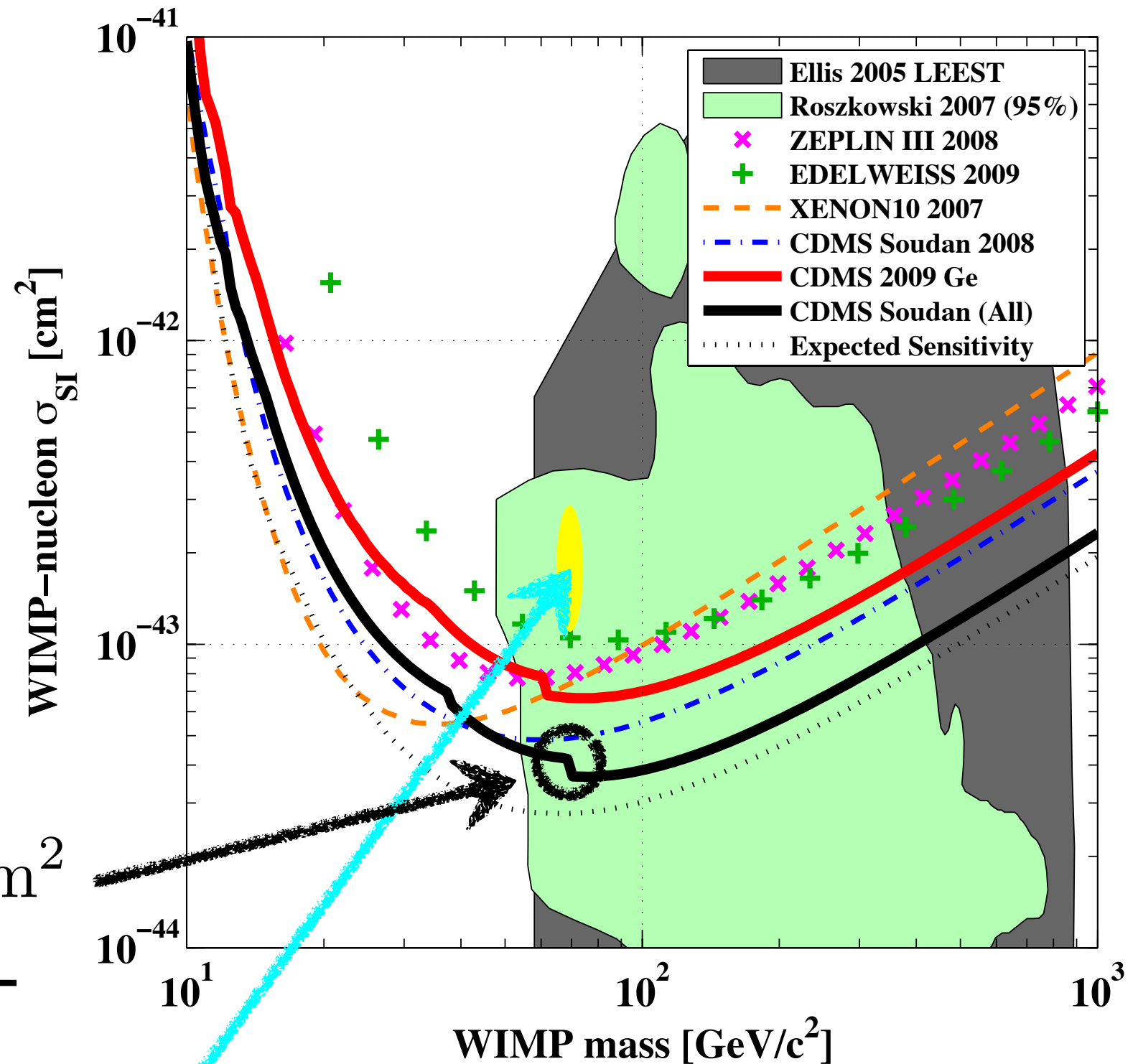
Exp. bound:

$$\sigma_{\text{SI}} \lesssim 3.8 \times 10^{-44} \text{ cm}^2$$

90% CL

Dark Higgs

Prediction: $\sigma_{\text{SI}} \simeq (1.2 - 2.7) \times 10^{-43} \text{ cm}^2$



Uncertainties in the direct detection

Local density of CDM (not measured)

$$\rho_0 = 0.3 \text{ GeV}/\text{cm}^3$$

assumed in the experiments.

$$\rho_0 = 0.2 \sim 0.6 \text{ GeV}/\text{cm}^3$$

reasonable for smooth halo.

$$\rho_0 \sim 0.04 \text{ GeV}/\text{cm}^3 \text{ (Kamionkowski and Koushiappas)}$$

possible for non-smooth halo.

Effective Higgs coupling $HH\bar{f}f$

may be altered in more general models.

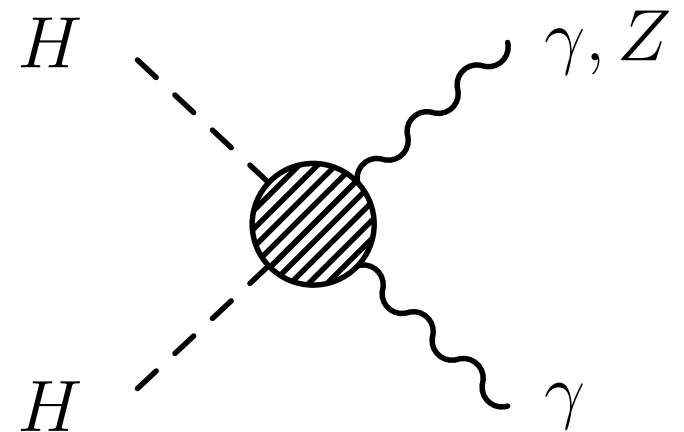
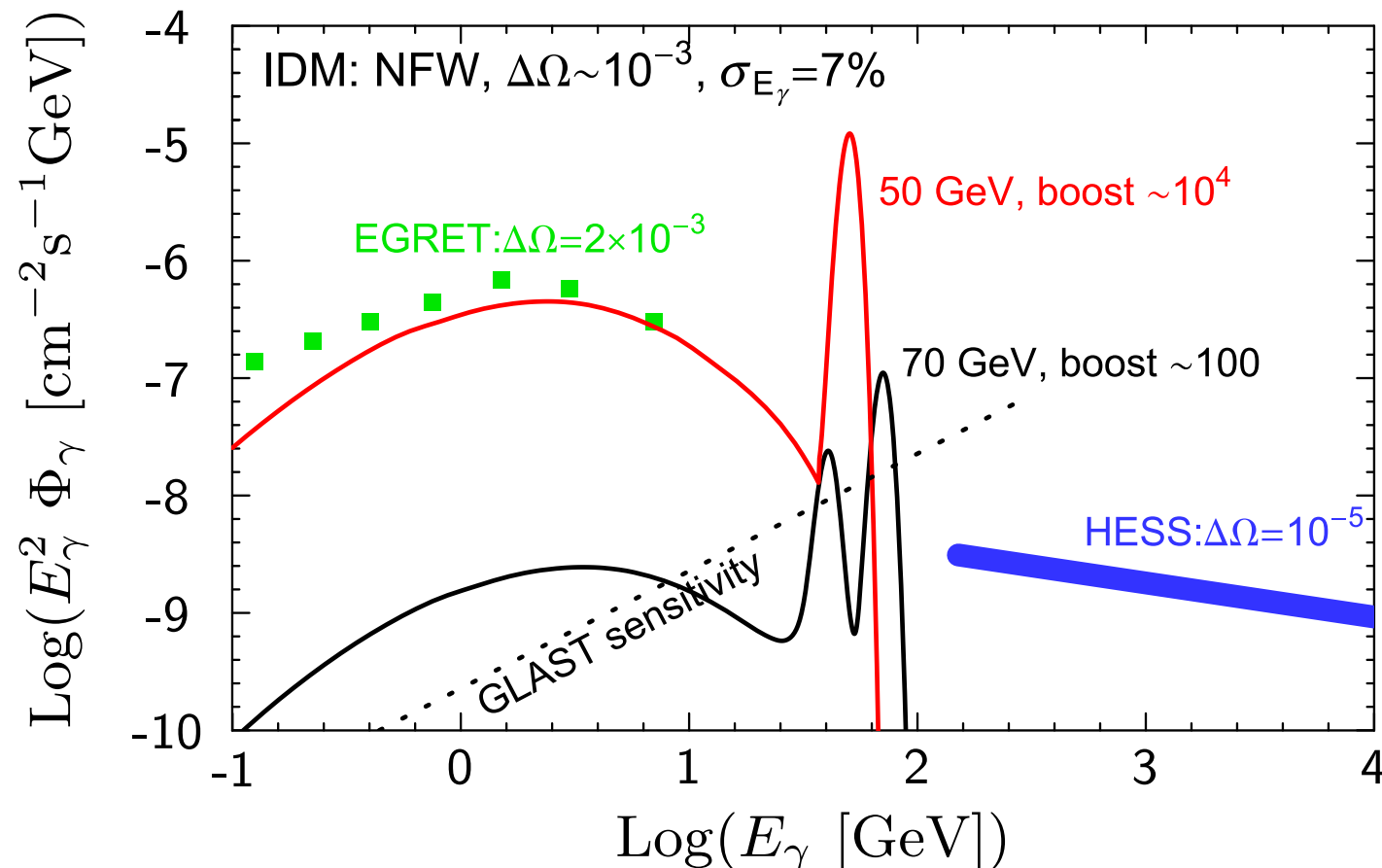
Astrophysical Signals

$HH \rightarrow \gamma\gamma, \gamma Z$ in the Galactic halo.

Two (nearly) monochromatic gamma lines.

$$E_\gamma = m_H (\simeq 70\text{GeV}), m_H - m_Z^2/(4m_H) (\simeq 40\text{GeV})$$

$$\sigma_{\gamma\gamma(\gamma Z)} v|_{v \rightarrow 0} \simeq 4.3(5.4) \times 10^{-29} \text{cm}^3/\text{s}$$



cf. Inert Doublet Model
 ← Gustafsson et al.